

**Satellite Communication and Positioning Systems – Prof. L. Luini,  
June 27<sup>th</sup>, 2023**

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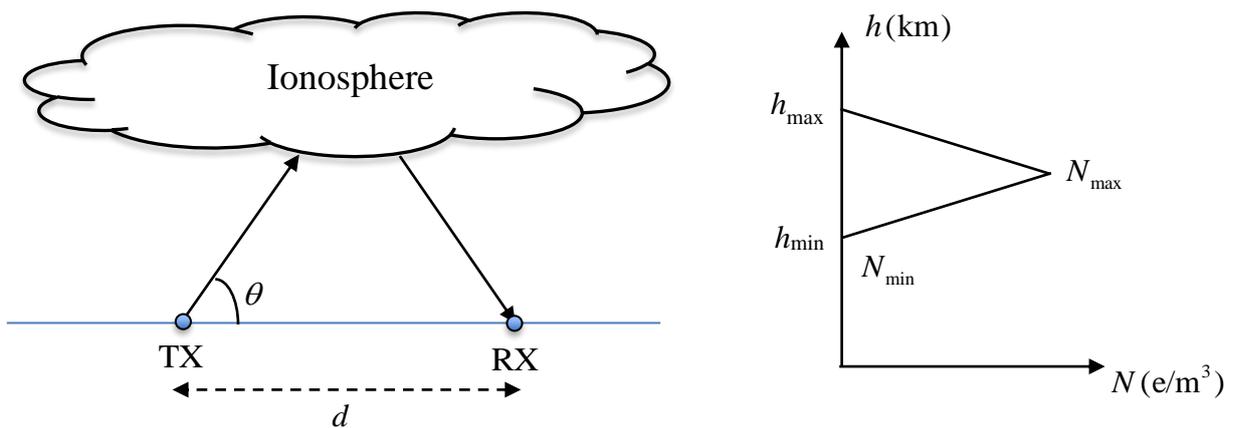
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**Problem 1**

Making reference to the figure below, we want the transmitter TX to reach the user RX at distance  $d$  by exploiting the ionosphere (elevation angle  $\theta = 60^\circ$ ). The ionosphere is modelled with the symmetric electron density profile (daytime) sketched in the figure (right side), where  $N_{\max} = 6 \times 10^{12} \text{ e/m}^3$ ,  $N_{\min} = 4 \times 10^{10} \text{ e/m}^3$ ,  $h_{\min} = 100 \text{ km}$  and  $h_{\max} = 400 \text{ km}$ .

- 1) Determine the maximum distance  $d$  achievable for the TX  $\rightarrow$  RX link.
- 2) Determine the operational frequency  $f$  to achieve the conditions at point 1).
- 3) Indicate a reasonable margin on  $f$  found at point 2) to guarantee the TX  $\rightarrow$  RX link notwithstanding ionospheric variations.
- 4) Indicate the best polarization to be used for the TX  $\rightarrow$  RX link.

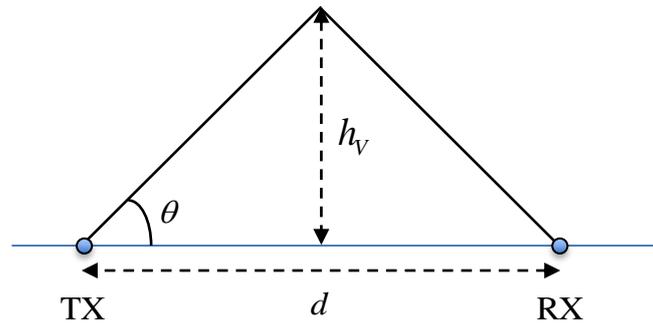
Assume: the virtual reflection height  $h_V$  is 1.2 times  $h_R$ , the height at which the wave is actually reflected; the Earth is flat; no tropospheric effects to be considered.



**Solution**

1) The distance  $d$  is maximized if the reflection occurs as high as possible in the ionosphere, i.e. at the height  $h_p = 250 \text{ km}$ , correspondent to  $N_{\max}$ . Considering the figure below, the distance can be found by inverting the following expression:

$$h_v = 1.2 h_p = d/2 \tan\theta \rightarrow d = \frac{2.4 h_p}{\tan\theta} = 346.4 \text{ km}$$



2) The link operational frequency  $f$  can be determined by inverting the following equation:

$$\cos\theta = \sqrt{1 - \left(\frac{f_c}{f_m}\right)^2} = \sqrt{1 - \left(\frac{9\sqrt{N_{\max}}}{f_m}\right)^2}$$

Solving for the frequency  $f_m$ , we obtain:

$$f_m = \sqrt{\frac{81N_{\max}}{1 - [\cos(\theta)]^2}} = 25.5 \text{ MHz}$$

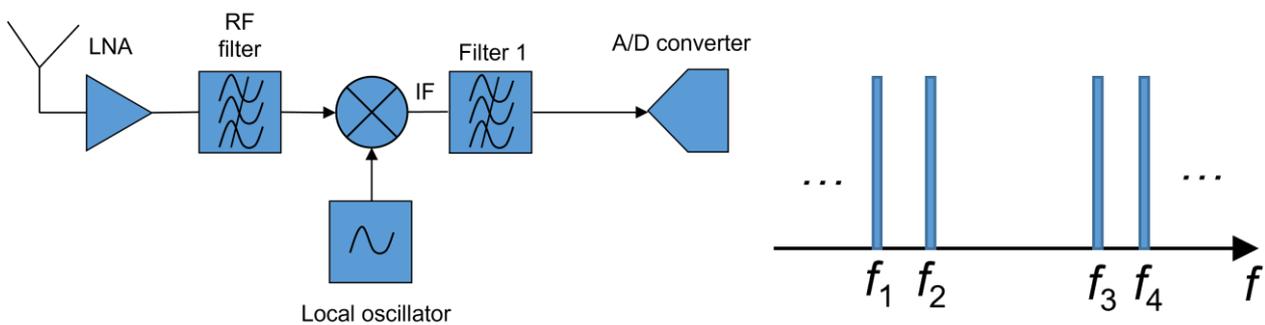
3) During the night, the peak values of the electron content will decrease: it is a good rule of thumb to use reduce by 10% the peak frequency to avoid that the wave crosses the ionosphere at nighttime. Therefore  $\rightarrow f' = 0.9f = 22.95 \text{ MHz}$ .

4) Depolarization in the ionosphere affects linearly polarized waves, but not circularly polarized ones. Therefore, the best polarization is LHCP or RHCP.

## Problem 2

Consider the heterodyne receiver depicted below (left side), which aims at receiving the RF signal with carrier frequency  $f_2 = 20$  GHz, to be digitalized at the end of the receiver chain by using an A/D converter with a maximum sampling frequency  $f_s = 4.2$  GS/s. As indicated in the picture below (right side), the RF spectrum is occupied by multiple signals, all having the same bandwidth  $B = 200$  MHz, with the closest carriers being  $f_1 = 19$  GHz and  $f_3 = 25$  GHz (more signals are present as indicated in the figure).

- 1) Design the RF filter and local oscillator considering the A/D converter features and the aim of reducing as much as possible the RF filter complexity (i.e. selectivity): specifically, provide a suitable  $f_{LO}$ , indicate the type of filter to be used for the RF filter and its cutoff(s) frequency(ies).
- 2) Indicate the ideal bandwidth of Filter 1.



## Solution

1) When down-converting signals from RF to intermediate frequency (IF), image signals represent a problem. The same IF can be obtained using a local oscillator  $f_{LO}$  higher or lower than the target carrier frequency  $f_2$ . If  $f_{LO} < f_2$ , the image signals to be rejected using the RF filter are those lower than  $f_2$ ; if  $f_{LO} > f_2$ , the image signals to be rejected using the RF filter are those higher than  $f_2$ . As a consequence, the RF filter for the case  $f_{LO} < f_2$  will need to be much more selective than the one for the  $f_{LO} > f_2$  case. Therefore, to meet the requirements expressed at point 1, the  $f_{LO} > f_2$  case is to be selected. To reduce the RF filter complexity (low-pass filter) as much as possible, the ideal cutoff frequency is, for example, 22.5 GHz, exactly in the middle of the  $f_3$ - $f_2$  interval (other choices are possible). Regarding the value for the local oscillator frequency, if  $f_{LO} = 2.5$  GHz, then  $f_{IF} = f_{LO} - f_2 = 2.5$  GHz, with a maximum frequency of the IF signal equal to  $f_{max} = 2.6$  GHz. However, according to the Nyquist theorem, in this case, the minimum sampling frequency of the A/D converter should be  $f_s = 2f_{max} = 5.2$  GS/s. This exceeds the available specifications for the A/D converter, i.e.  $f_s = 4.2$  GS/s. Using this value as an additional constraint, a possible final optimum design is:  $f_{LO} = 22$  GHz,  $f_{IF} = f_{LO} - f_2 = 2$  GHz,  $f_{max} = 2.1$  GHz,  $f_s = 2f_{max} = 4.2$  GS/s.

2) Given the design at point 1, the optimum Filter 1 (bandpass filter) will have a lower cut-off frequency of  $f_{min} = f_{IF} - B/2 = 1.9$  GHz and  $f_{max} = f_{IF} + B/2 = 2.1$  GHz.

### Problem 3

We need to design a link to a deep-space probe orbiting Mars and operating at Ka-band, specifically at  $f = 26$  GHz. The ground station is equipped with a steerable antenna to track the probe.

- 1) Calculate the reflector antenna diameter of the ground station (Gregorian configuration with efficiency  $\eta = 0.5$ ) necessary to guarantee that the probe can be correctly tracked down to an elevation angle  $\theta = 30^\circ$  for 99.9% of the time in a year, i.e. that the minimum SNR is 5 dB.

To this aim, assume:

- that the atmosphere is stratified;
- that the ground station LNA noise temperature is  $T_R = 50$  K;
- to neglect the cosmic background temperature;
- that the mean radiating temperature of the atmosphere is  $T_{mr} = 10$  °C;
- that the probe makes use of a parabolic antenna with gain  $G_T = 45$  dB;
- the transmit power is  $P_T = 110$  W;
- the probe antenna always points at the ground station;
- the distance between the probe satellite and the ground station is  $L = 225000000$  km;
- the receiver bandwidth is  $B = 1$  kHz;
- that the CCDF of the zenithal atmospheric attenuation  $A_T$  is modelled by:

$$P(A_Z^{dB}) = 100e^{-0.69A_Z^{dB}} \quad (A_Z \text{ in dB and } P \text{ in } \%)$$

- 2) Calculate the maximum data rate achievable for this channel with the conditions at point 1), considering a negligible bit error rate.

### Solution

1) The zenithal attenuation  $A_Z^{dB}$  is determined using the CCDF model. 99.9% availability corresponds to  $P = 0.1\%$  exceedance. Inverting the CCDF formula:

$$A_Z^{dB} = -\frac{1}{0.69} \ln\left(\frac{0.1}{100}\right) \approx 10 \text{ dB}$$

Scaling the zenithal attenuation to the target elevation angle:

$$A_S^{dB} = \frac{A_Z^{dB}}{\sin(\theta)} \approx 20 \text{ dB}$$

which, in linear scale, corresponds to:

$$A_L = 10^{\frac{A_S^{dB}}{10}} \approx 0.01$$

The system noise temperature is (for the Gregorian configuration, the waveguide is very short and its effect on the noise can be neglected):

$$T_{sys} = T_R + T_A = T_R + T_{mr}(1 - A_L) = 330.3 \text{ K}$$

The SNR is given by:

$$SNR = \frac{P_T G_T f_T (\lambda/4\pi L)^2 G_R f_R A_L}{k T_{sys} B}$$

where  $f_R = 1$  and  $f_T = 1$ .

Inverting the expression above to solve for  $G_R$  (by setting  $SNR = SNR_{min} = 5$  dB):

$$G_R \approx 74 \text{ dB}$$

Recalling that:

$$\frac{\eta A_g}{G_R} = \frac{\lambda^2}{4\pi}$$

where  $A_g$  is the geometrical area of the antenna:

$$A_g = \left(\frac{D_R}{2}\right)^2 \pi$$

the antenna diameter  $D_R$  is obtained:

$$D_R \approx 26 \text{ m}$$

This is indeed the dimension of Ka-band deep-space antennas installed at NASA Deep Space Network (DSN) sites (Goldstone, Madrid, Canberra).

2) The reply is given by the Shannon limit:

$$C = B \log_2(1 + SNR_{min}) \approx 2 \text{ kb/s}$$

### Problem 4

A two-frequency GNSS receiver, installed at sea level, is correctly providing the PVT solution. The equivalent noise temperature of the receiver is  $T_R = 1000$  K. The error due to the code correlator is associated to the SNR after despreading as follows (consider the L1 C/A code):

$$d^{C/A} = 60SNR^{-0.2} \quad (\text{SNR in dB, } d^{C/A} \text{ in m})$$

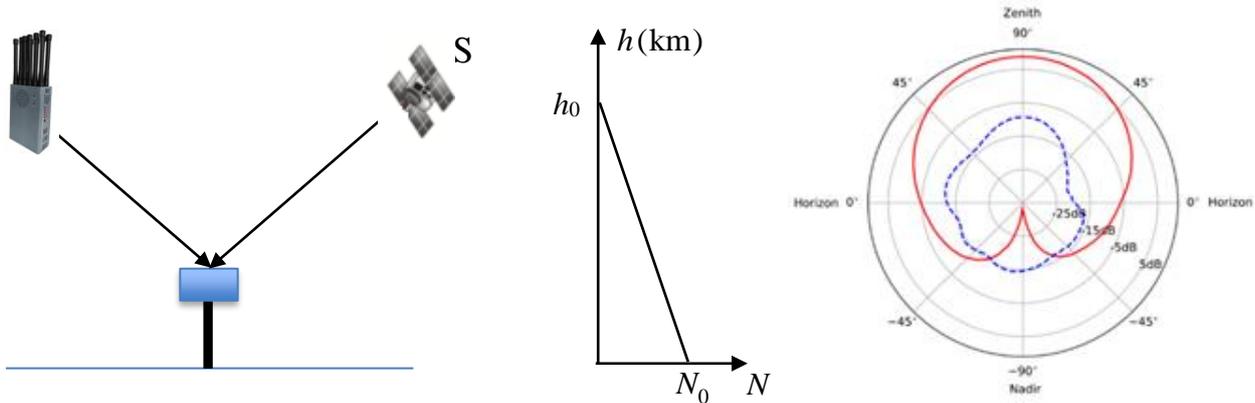
The refractivity profile is modeled as sketched in the figure below, where  $N_0 = 900$  ppm and  $h_0 = 9.5$  km, while the zenithal TEC value is 40 TECU.

- 1) Determine the range error for a specific GPS satellite, visible at elevation angle  $\theta = 45^\circ$ , at  $L = 20000$  km from the receiver.

OPTIONAL: at some point, a jammer, also at  $45^\circ$  elevation angle, begins transmitting a wideband signal evenly covering the full L band and reaching the GPS receiver antenna with power spectral density  $S_J$  in  $\text{W}/(\text{m}^2 \cdot \text{Hz})$ .

- 2) For the same satellite, determine the maximum value of  $S_J$  for the range error to be limited below 42 m.

Assume that: the zenithal atmospheric attenuation is  $A_Z = 0.4$  dB, the mean radiating temperature is  $T_{mr} = 260$  K, the distance between the receivers and the satellites is  $L = 20000$  km.



### Solution

As the GNSS receiver is correctly providing the PVT solution, there is no clock bias. The pseudorange is therefore affected by the following error sources:

$$\rho = r + d^{C/A} + d^I + d^T$$

where:

$$d^T = \frac{h_0 N_0 10^{-6}}{2 \sin(\theta)} = 6.05 \text{ m is the slant tropospheric error (given the profile in the figure).}$$

$$d^I \approx 0 \text{ m, as a dual frequency receiver is considered.}$$

$d^{C/A}$  depends on the SNR.

The L1 frequency is  $f = 1575.42$  MHz, so the wavelength is  $\lambda = c/f = 0.1904$  m. The received power can be calculated as:

$$P_R = P_T G_T f_T \left( \frac{\lambda}{4\pi L} \right)^2 G_R f_R A_{ATM}$$

Making reference to the specifications of GPS satellites, we can assume  $P_T = 21.9$  W and  $G_T f_T = 13.4$  dB = 21.9 (worst case). The slant path atmospheric attenuation is:

$$A = A_Z / \sin(45^\circ) = 0.57 \text{ dB} \rightarrow A_{lin} = 0.88$$

The gain of the receiver antenna can be derived from the radiation pattern figure: at  $45^\circ$ , the receiver gain is  $G_R = 5 \text{ dB} = 3.16$ . This value also includes  $f_R$ .

The received power is calculated directly from the equation above, yielding:

$$P_R = 7.63 \times 10^{-16} \text{ W}$$

The system noise power depends on the antenna noise temperature  $T_A$  and on the receiver noise temperature  $T_R$ :

$$T^{sys} \approx T_C A_{lin} + T_{mr}(1 - A_{lin}) + T_R = 1034.1 \text{ K}$$

After despreading, the bandwidth for the calculation of the noise power is  $B = 50 \text{ Hz}$ . The noise power is calculated as:

$$P_N = kT_{sys}B = 7.14 \times 10^{-19} \text{ W}$$

As a result:

$$SNR = 30.29 \text{ dB} \rightarrow d^{C/A} = 30.3 \text{ m.}$$

$$\text{The total range error is: } d = d^{C/A} + d^T = 36.4 \text{ m.}$$

2) The jammer will increase the noise power as follows:

$$P'_N = P_N + S_J B A_E = P_N + S_J B G_T f_T \frac{\lambda^2}{4\pi}$$

Considering that  $d^T$  is fixed, and setting  $d = 42 \text{ m} \rightarrow d^{C/A} = 35.95 \text{ m} \rightarrow SNR_m = 12.94 \text{ dB} = 19.7$ .

As a result:

$$P'_N = \frac{P_R}{SNR_m} = 3.88 \times 10^{-17} \text{ W}$$

Therefore:

$$S_J B G_T f_T \frac{\lambda^2}{4\pi} = P'_N - P_N = 3.8 \times 10^{-17} \text{ W} \rightarrow S_J = 1.2 \times 10^{-17} \frac{\text{W}}{\text{m}^2 \text{Hz}}.$$